





# The SWFO Program and its Science Objectives

Space Weather Follow On – Lagrange 1 (SWFO-L1) Pre-Launch Science Meeting September 21, 2025

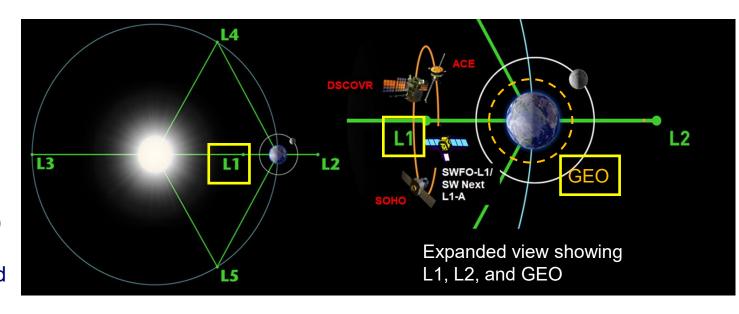
Dimitris Vassiliadis NOAA/NESDIS/SWO

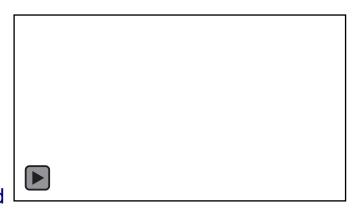


### The Space Weather Follow On Program



- SWFO is designed to provide critical solar and heliospheric data to users: both to operational users (forecasters, satellite operators, others), and retrospective users (scientists, other developers).
- SWFO will be operating from two orbits:
  - GEO: The Compact Coronagraph-1 (CCOR-1) on the GOES-19 satellite (launched on June 25, 2024)
  - Sun-Earth Lagrange 1 (L1) point: The SWFO-Lagrange 1 (SWFO-L1) Observatory with a planned launch of <u>September 23, 2025</u> as a rideshare with IMAP.
- SWFO will be providing the following products:
  - Coronal imagery to enable early situational awareness for forecasting
  - Solar wind and interplanetary magnetic field (IMF) measurements to be used as inputs to geospace models.
  - Suprathermal particle flux measurements to be used to improve estimates of CME arrival times.



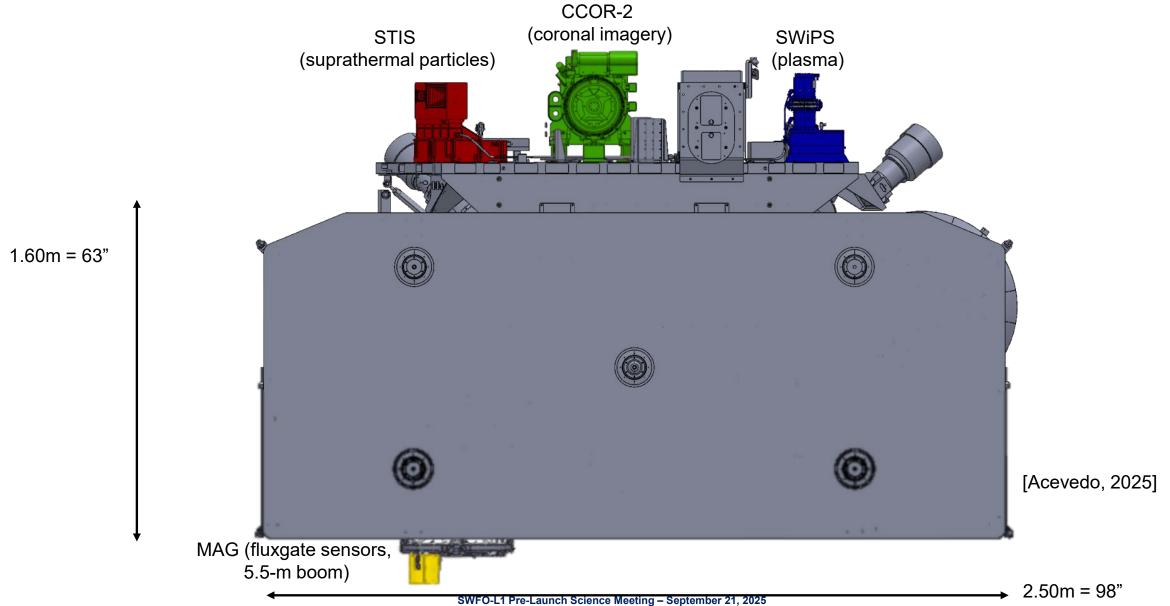






#### **SWFO-L1 Instrumentation**





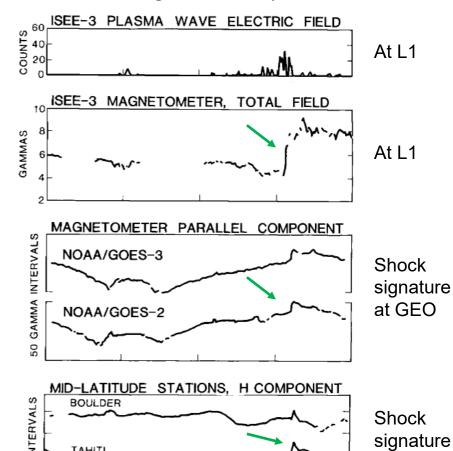


### Solar Wind L1 Monitors: a NASA-NOAA Collaboration from More than 45 Years Ago



- SWFO-L1 is part of a historic line of real-time solar wind monitors at L1 that started with a NASA mission in the 1980s.
- Excerpt from "ISEE 3 in real time: an update" by Joselyn et al. of the Space Environment Laboratory (now SWPC) [AGU Eos 62, 32, Aug 1981]:
  - In the October 9, 1979 issue of Eos, Tsurutani and Baker [1979] expressed the tangible benefits of obtaining selected interplanetary data from the ISEE 3 spacecraft in real time.
  - ... ISEE 3 is positioned between the earth and the sun in a halo orbit about the sun-Earth libration point. This location is well suited to provide advanced warning of the onset of geomagnetic disturbances. Although the data were not originally intended for real-time use, NOAA and NASA began to cooperate early in 1979 to solve the technical and administrative problems requisite to acquiring the raw data before editing and transmission to the experimenters. In March 1980 the data stream began arriving at the Space Environment Services Center (SESC), Boulder, Colorado. Daily data coverage varies but averages near 80%. The data are now being routinely used to support SESC military and civilian customers and the scientific community at large.
- The ISEE 3 real-time data provision ended in October 1982, but eventually led to the concept of ACE as a solar wind monitor.

Measuring an interplanetary shock at L1 and resulting SpWx effects at GEO and on the ground [Joselyn et al., 1981]



UNIVERSAL TIME

HONOLULU

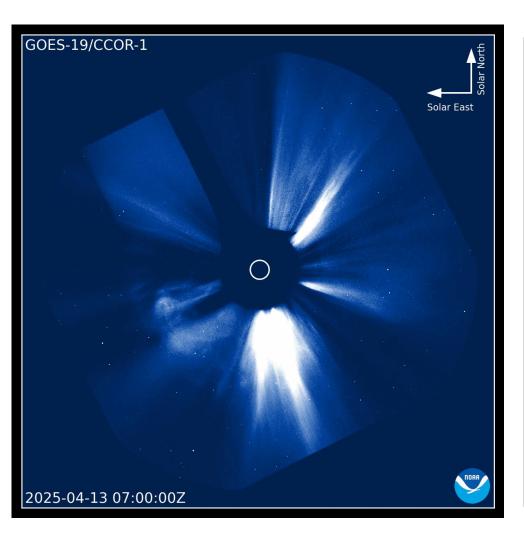
on the

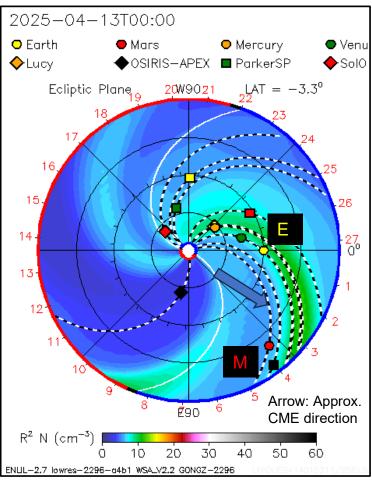
ground



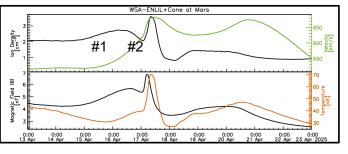
### Integration of CCOR-1 Coronal Imagery in SWPC and M2M Operations: Halo CMEs of April 13, 2025



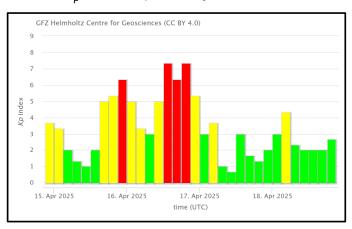








Earth: G4 geomagnetic storm observed. Below:  $K_p$  index of planetary-level disturbance

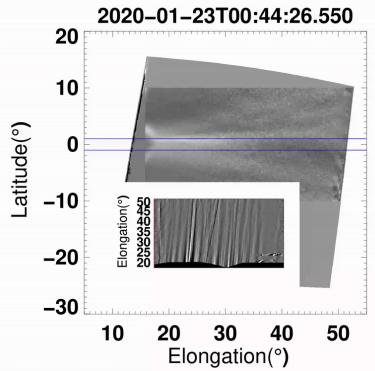




#### **CCOR Imagery: Feature Tracking Applications for Modeling and Theory**



- CCOR-1 (and soon, CCOR-2) images are summarily used to detect large-scale CME features. However, their applications may be extended to measure the 1D or 2D velocity fields of much smaller density structures. The resulting velocity maps will be useful for informing high-resolution models of outer corona.
- Examples from the recent literature are shown.



Flow tracking based on PSP heliospheric images [Nindos et al., A&A 2020]

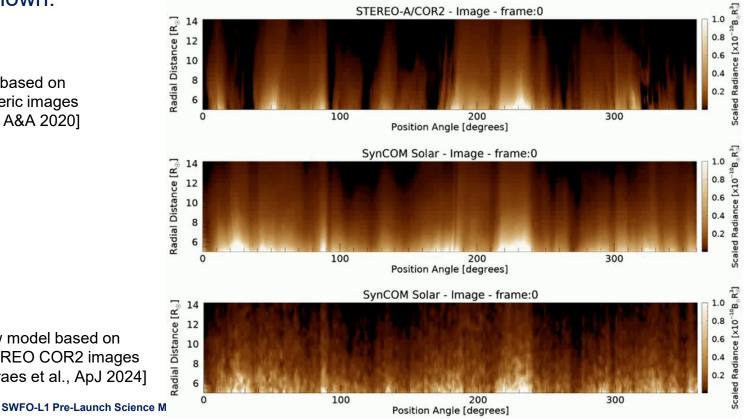
> Flow model based on STEREO COR2 images [Moraes et al., ApJ 2024]



Is there a boundary between the corona and the solar wind?

Research challenge: identifying the corona-solar wind boundary

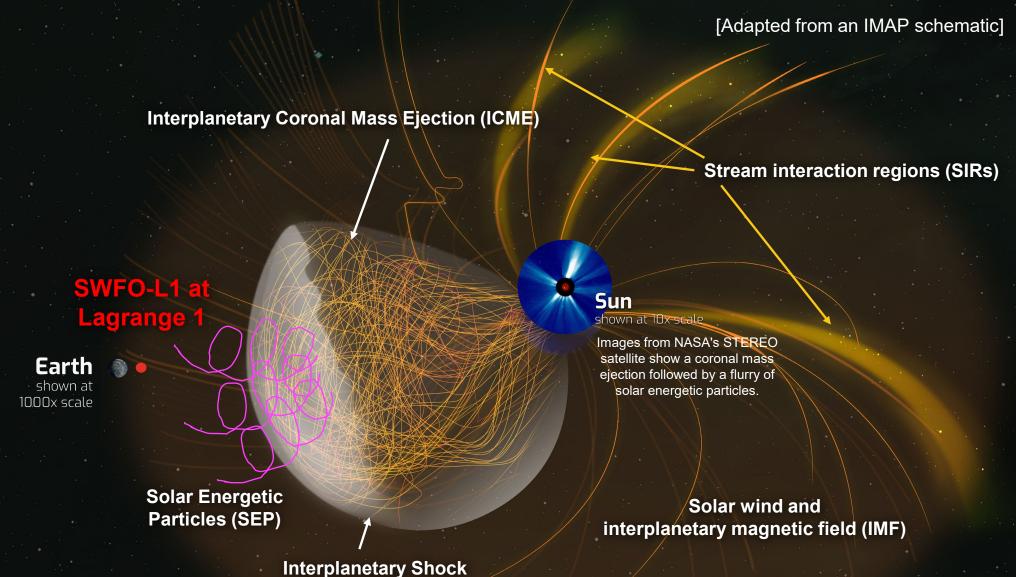
[Cranmer, SHINE 2025; background image: M. Druckmueller1





### In Situ Measurements: Probing the Internal Structure of ICMEs



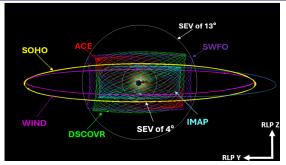


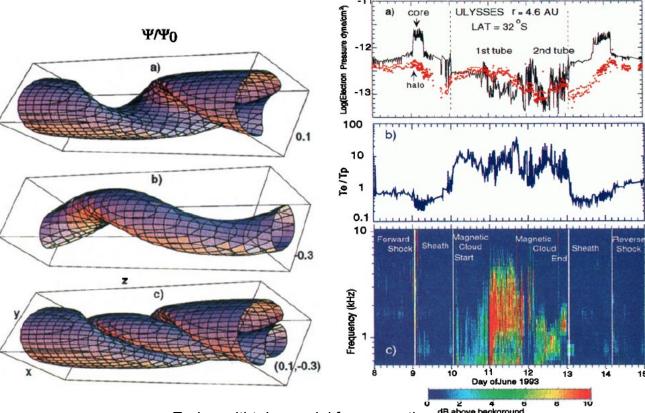


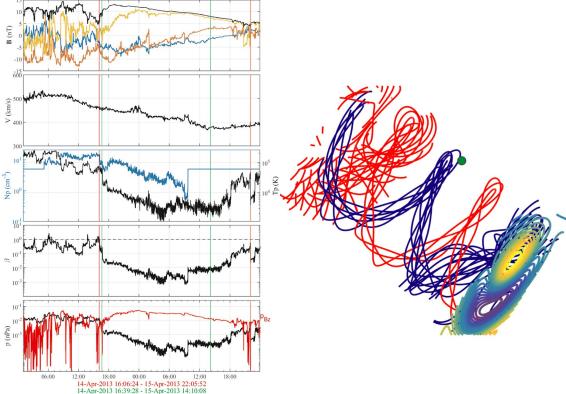
## In Situ Measurements: Probing the Internal Structure of ICMEs



 Magnetic clouds within ICMEs are often modeled as a multiscale helically symmetric system of flux tubes. Force-free models (Lundquist, Freidberg) are obtained from regressions to B-field and plasma data.  The availability of several solar wind monitors at L1 will enable a detailed look at the magnetic and thermodynamic properties of these equilibria.







Early multi-tube model for magnetic cloud [Osherovich et al., GRL 1999]

Recent model applied to ACE data to obtain a magnetic cloud configuration [Hu et al., GRL 2021]



### **Summary**



- The SWFO Program is designed provide continuity in coronal imagery and solar wind L1
  measurements for NOAA's Space Weather Prediction Center. Its data products will have dual
  use in supporting situational awareness and furthering solar and heliospheric research. The
  CCOR-1 coronagraph launched as part of the GOES-19 satellite on June 25, 2024 has been
  recording coronal images since September 19, 2024, which have been validated and released
  publicly since February 22, 2025.
- The SWFO-L1 observatory is planned to be launched from Kennedy Space Center on September 23, 2025 as a rideshare with IMAP and Carruthers.
- Based on the SWFO-L1 operation from a Lagrange 1 orbit, its data are expected to be used in a wide range of geospace applications including the driving of real-time models of the space environment and supporting model development and analysis. The measurements will provide a key reference for observations from missions in other heliospheric locations.

Go SWFO-L1!